

## Strain history of ice shells of the Galilean satellites from radar detection of crystal orientation fabric

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Received 29 December 2010; revised 4 February 2011; accepted 11 February 2011; published 24 March 2011.

[1] Orbital radar sounding has been suggested as a means of determining the subsurface thermal and physical structure of the outer ice I shells of the Galilean satellites. At radar frequencies, the dielectric permittivity of single- and polycrystalline water ice I is anisotropic. Crystal orientation fabric (COF), which is indicative of strain history, can be unambiguously detected by comparing the received power of dual co-polarization (linear polarization parallel and perpendicular to the orbit) radar data. Regions with crystal orientations dictated by the local strain field (“fabric”) form in terrestrial ice masses where accumulated strain and temperature are high, similar to conditions expected in a convecting outer ice I shell on Europa, Ganymede, or Callisto. We use simulations of solid-state ice shell convection to show that crystal orientation fabric can form in the warm convecting sublayer of the ice shells for plausible grain sizes. Changes in received power from parallel and perpendicular polarizations in the ice shells due to fabric could be detected if multi-polarization data is collected. With proper instrument design, radar sounding could be used to shed light on the strain history of the satellites’ ice shells in addition to their present day internal structures. **Citation:** Barr, A. C., and D. E. Stillman (2011), Strain history of ice shells of the Galilean satellites from radar detection of crystal orientation fabric, *Geophys. Res. Lett.*, 38, L06203, doi:10.1029/2010GL046616.

### 1. Introduction

[2] The next mission to the Jovian satellites will likely carry an ice-penetrating radar to study the subsurface physical and thermal structure of the floating outer ice I shells of Europa and Ganymede [Blankenship *et al.*, 2009]. Because radar attenuation in ice is caused primarily by temperature and salinity [Evans, 1965; Glen and Paren, 1975], studies of the ice shells’ internal temperature and salt distributions [Chyba *et al.*, 1998; Moore, 2000] have been key for constraining sounder design parameters.

[3] The orientation of the ice crystal lattice also affects the permittivity of ice. Deformation in polycrystalline ice can change the size and orientation of the crystal lattice of individual grains [Alley, 1988]. When a stress is applied, grains deform by basal slip and their crystallographic *c* axes rotate toward the axis of compression [Alley, 1988]. Large strains at high temperatures create regions wherein the lattices of adjacent crystals are co-aligned (crystal orientation fabric or COF). Ice core samples suggest that fabric forms

where the total accumulated strain,  $\epsilon > 0.5$  [Alley, 1988; Budd and Jacka, 1989]. Seasonally average temperatures in most ice cores are  $T > 238$  K ( $-35^{\circ}\text{C}$ ). These conditions are similar to conditions within the warm regions of a convecting ice shell on the Galilean satellites [Tobie *et al.*, 2003].

[4] On Earth, COF detected by seismic [Blankenship, 1989] and radar sounding [Siebert and Kwok, 2000; Siebert and Fujita, 2001; Matsuoka *et al.*, 2003; Dall, 2009, 2010] has been used to infer the strain history of ice sheets. For single-crystal and polycrystalline ice with well-developed COF ( $\epsilon \gg 1$  at high temperature), the difference between the real part of the dielectric permittivity  $\epsilon$ , in a direction parallel to ( $\epsilon_{\parallel}$ ) and perpendicular to ( $\epsilon_{\perp}$ ) the *c*-axis is about 1% [Johari and Charette, 1975; Fujita *et al.*, 1993], which causes radar reflections at abrupt contrasts in COF, anisotropic reflectivity, and birefringence [Matsuoka *et al.*, 2003]. Reflections due to changes in COF can be as large as  $\sim 50$  dB [Fujita *et al.*, 1993], within the expected  $-60$  dB dynamic range for a proposed radar sounder for Europa [Blankenship *et al.*, 2009], but these reflections cannot be separated from reflections caused by changes in conductivity or density. Anisotropic reflectivity and birefringence vary periodically with polarization angle [Hargreaves, 1977; Matsuoka *et al.*, 2003]. Zones of anisotropic reflectivity and birefringence have been detected by comparing the received power of dual co-polarization (i.e., linear polarization parallel and perpendicular to the ground track) ice-penetrating radar data [Matsuoka *et al.*, 2003; Dall, 2009, 2010]. Polarization-dependent reflectivity provides an unambiguous means of detecting COF and could be used to test hypotheses about ice shell heat transfer and modes of resurfacing.

[5] Here, we simulate solid-state convection in the outer ice I shells of Europa, Ganymede, and Callisto to determine whether the strain and temperature conditions are conducive to COF formation. A composite rheology [van den Berg *et al.*, 1993; McNamara *et al.*, 2001; Freeman *et al.*, 2006] for water ice I based on laboratory experiments and microphysical models of the relevant deformation processes [Goldsby and Kohlstedt, 2001] is used. COF forms in locations where convective strain is accommodated principally by anisotropic creep mechanisms dislocation and grain size-sensitive creep [Alley, 1988; Goldsby and Kohlstedt, 2002; Karato, 2008]. We show that COF can form in the warm convecting sublayer of an ice I shell for  $d > 1$  mm. If the ice shells have a smaller uniform grain size, fabric forms only in convective downwellings. We also discuss implications for radar sounder design and the formation of COF beneath surface features such as Europa’s bands and Ganymede’s grooved terrain.

### 2. The Model

[6] Laboratory experiments suggest that the total strain rate ( $\dot{\epsilon}$ ) in ice is equal to the sum of strain rates from

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dislocation creep, grain size-sensitive creep (which occurs by grain boundary sliding and basal slip), and diffusion creep [Goldsby and Kohlstedt, 2001],

$$\dot{\epsilon} = \dot{\epsilon}_{disl} + \left( \dot{\epsilon}_{gbs}^{-1} + \dot{\epsilon}_{bs}^{-1} \right)^{-1} + \dot{\epsilon}_{diff}, \quad (1)$$

where the strain rate from each deformation mechanism is related to temperature ( $T$ ), stress ( $\sigma$ ), and grain size ( $d$ ),

$$\eta_i = \frac{d^{p_i}}{2A_i} \sigma^{1-n_i} \exp\left(\frac{Q_i^*}{RGT}\right), \quad (2)$$

with  $\eta = \sigma/2\dot{\epsilon}$ , grain size exponent (for rheology  $i$ )  $p_i$ , stress exponent  $n_i$ , pre-exponential constant  $A_i$ , gas constant  $R_G = 8.314 \text{ J mol}^{-1} \text{ K}^{-1}$  and activation energy  $Q_i^*$ . For dislocation creep,  $A_{disl} = 4.0 \times 10^{-19} \text{ Pa}^{-4} \text{ s}^{-1}$ ,  $n_{disl} = 4$ ,  $p_{disl} = 0$ , and  $Q_{disl}^* = 60 \text{ kJ mol}^{-1}$ ; for grain boundary sliding,  $A_{gbs} = 6.2 \times 10^{-14} \text{ m}^{1.4} \text{ Pa}^{-1.8} \text{ s}^{-1}$ ,  $n_{gbs} = 1.8$ ,  $p_{gbs} = 1.4$ ,  $Q_{gbs}^* = 49 \text{ kJ mol}^{-1}$ ; and for basal slip,  $A_{bs} = 2.2 \times 10^{-7} \text{ Pa}^{-2.4} \text{ s}^{-1}$ ,  $n_{bs} = 2.4$ , and  $p_{bs} = 0$ , and  $Q_{bs}^* = 60 \text{ kJ mol}^{-1}$ . The viscosity for diffusion creep is [Goodman et al., 1981; Barr and Showman, 2009],

$$\eta_{diff} = \frac{1}{2} \frac{3RGT_b d^2}{42V_m D_{o,v}} \exp\left(\frac{Q_v^*}{RGT}\right), \quad (3)$$

with molar volume  $V_m = 1.97 \times 10^{-5} \text{ m}^{-3}$ , volume diffusion coefficient  $D_{o,v} = 9.10 \times 10^{-4} \text{ m}^2 \text{ s}^{-1}$ , activation energy  $Q_v^* = 59.4 \text{ kJ mol}^{-1}$ , and basal temperature of the shell  $T_b$ . A composite rheology [van den Berg et al., 1993; McNamara et al., 2001; Freeman et al., 2006],

$$\frac{1}{\eta_{tot}} = \frac{1}{\eta_{disl}} + [\eta_{GSS} + \eta_{BS}]^{-1} + \frac{1}{\eta_{diff}}, \quad (4)$$

is used to account for contributions to the ice viscosity from each deformation mechanism.

[7] The same processes that drive  $c$ -axis orientation can form new grains by recrystallization [Alley, 1988], which can be tens to hundreds of mm in the Galilean satellites [Barr and McKinnon, 2007]. However, in terrestrial ice sheets with  $\dot{\epsilon}$  and  $T$  similar to a convecting ice shell,  $c$ -axis rotation is thought to occur without concomitant recrystallization [Alley, 1988]. Accordingly, we assume that the ice shell has a uniform grain size  $0.1 \text{ mm} < d < 30 \text{ mm}$  kept constant by grain boundary pinning from small silicate particles [Kirk and Stevenson, 1987; Barr and McKinnon, 2007].

[8] The finite-element model CITCOM [Moresi and Solomatov, 1995] is used to simulate convection for the range of  $d$  and ice shell thickness,  $D$ , appropriate for the satellites in a  $1 \times 1$  2-dimensional Cartesian box with  $64 \times 64$  elements. Three-dimensional Cartesian calculations in a  $4 \times 4 \times 1$  box with  $97 \times 97 \times 33$  elements are performed to study the pattern of COF-forming regions.

[9] The surface is kept at a constant  $T_s = 110 \text{ K}$ , appropriate for the mid-latitudes of Europa, Ganymede, and Callisto [Ojakangas and Stevenson, 1989; Herrick and Stevenson, 1990] and the bottom is heated to maintain  $T_b = 260 \text{ K}$ . Free-slip boundary conditions are

used on the sides of the domain. The Rayleigh number is,

$$Ra_o = \frac{\rho g \alpha \Delta T D^3}{\kappa \eta_o} \quad (5)$$

where  $\eta_o = \eta_{diff}(T = T_b)$  (see equation (3)),  $\Delta T = T_b - T_s$ , coefficient of thermal expansion  $\alpha = 10^{-4} \text{ K}^{-1}$ , ice density  $\rho = 920 \text{ kg m}^{-3}$ , thermal diffusivity  $\kappa = 10^{-6} \text{ m}^2 \text{ s}^{-1}$ , and gravity  $g = 1.3 \text{ m s}^{-2}$ , an average value roughly appropriate for all three satellites.

[10] For numerical tractability, we use a rheology of form [Solomatov and Moresi, 2000],

$$\eta_i = \frac{1}{\beta_i} \sigma^{1-n_i} \exp(E_i(1 - T')), \quad (6)$$

for each term in the composite flow law, where  $E_i = Q_i^* \Delta T / (R_G T_b^2)$ , and  $\beta_i$  describes the contribution of each component rheology [van den Berg et al., 1993],

$$\beta_i = (2\eta_o)^{n_i} \dot{\epsilon}_o^{n_i-1} A_i d^{-p_i} \exp\left(\frac{-Q_i^*}{R_G T_b}\right). \quad (7)$$

Sample calculations performed with an Arrhenius version of the temperature dependence (equation (2)) give similar temperature and strain fields in the warm region of the ice shell. In the composite flow law, the non-dimensional stress  $\sigma' = \sigma/\sigma_o$ ,  $\sigma_o = 2\eta_o \dot{\epsilon}_o$ , the reference strain rate  $\dot{\epsilon}_o = \kappa/D^2$ , and  $T' = (T - T_s)/(\Delta T)$ . Because the  $\beta$  factors depend on  $D$  and  $d$ , these properties must be specified *a priori*.

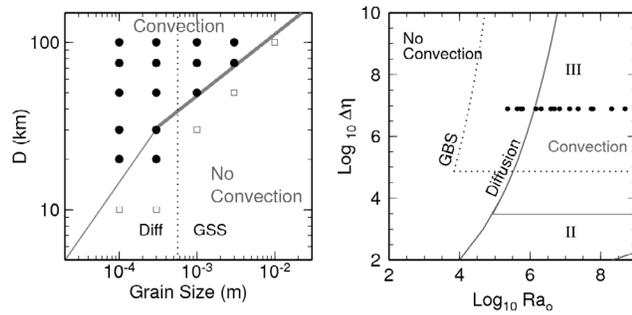
[11] We estimate strain in the ice shell as  $\epsilon \sim \dot{\epsilon}(D^2/\kappa)$ , by assuming that strain accumulates over at least one thermal diffusion time. Equation (6) underestimates the temperature dependence in the viscosity, so ice in the stagnant lid is unphysically soft and strains there should be interpreted as upper limits. COF forms where the ratio of strain due to dislocation + GSS creep to diffusion creep [McNamara et al., 2001],

$$R = \frac{\epsilon_{disl} + \epsilon_{GSS}}{\epsilon_{diff}}, \quad (8)$$

is greater than 1 (i.e., anisotropic strain dominates isotropic strain). Because little is known about COF formation at low temperatures, we also require  $T > 240 \text{ K}$ .

### 3. Results

[12] Figure 1 shows the values of  $D$  and  $d$  explored in the 2-dimensional Cartesian calculations. Figure 2 shows values of  $R$  in a convecting ice shell for several different uniform grain sizes. Regardless of grain size,  $R > 1$  in the stagnant lid, but temperatures there are too low to form COF. Beneath the stagnant lid,  $T > 240 \text{ K}$ ,  $R > 1$  and the total strain  $\gg 1$  in the center of the single convective roll if  $d > 1 \text{ mm}$ , a plausible grain size for the Galilean satellites' ice shells [Kirk and Stevenson, 1987; Barr and McKinnon, 2007]. Because radar attenuation in pure ice is temperature-dependent (attenuation increases by a factor of 10 between 248 and 272 K [Johari and Charette, 1975]), COF and interfaces between regions with different fabric orientations will be detected in cold downwellings and within the stagnant lid. If the stagnant lid is thin (i.e., the



**Figure 1.** (left) Values of grain size and ice shell thickness explored in 2-dimensional simulations. Gray lines show the locations of the critical ice shell thickness for convection from *Solomatov and Barr* [2007], circles indicate simulations where convection occurred, and squares show where convection did not occur. Dotted line indicates boundary between  $d$  values where GSS creep, rather than volume diffusion accommodates strain in the convective sublayer. (right) Parameter space in  $Ra_o - \Delta\eta$  space ( $\Delta\eta \equiv (\eta'_{tot}(T' = 0, \sigma' = 1))/(\eta'_{tot}(T' = 1, \sigma' = 1))$ ). Gray lines indicate boundaries between convection/no convection and the stagnant lid (III) and sluggish lid (II) regimes ( $\Delta\eta > \exp(4(n + 1))$  where  $n = 1$  for volume diffusion) [Solomatov, 1995]. Dotted line shows regime boundaries for GBS ( $n = 1.8$ ).

ice shell is thick and has a small grain size, a  $-50$  dB reflection would be detectable just above the stagnant lid. If the lid is thick (i.e., low  $Ra_o$  and large  $d$ ), only strong reflections,  $-10$  dB, could be detected at the base of the stagnant lid. In a salty ice shell (ice saturated with  $Cl^-$  [Grimm *et al.*, 2008]), the search for COF would be confined to old patterns preserved in the cold stagnant lid.

[13] Figure 3 illustrates the temperature and  $R$  values in a 30 kilometer-thick ice shell with  $d = 0.3$  mm. Values of  $R$  are highest in downwellings and lowest in the upwellings. A horizontal slice through the shell at the location where the average  $T = 240$  K shows that the diameter of regions with  $R > 1$  is about 25 km, similar to the thickness of the ice shell. An instrument that could spatially resolve convective upwellings would thus be able to detect fabric formed in downwellings.

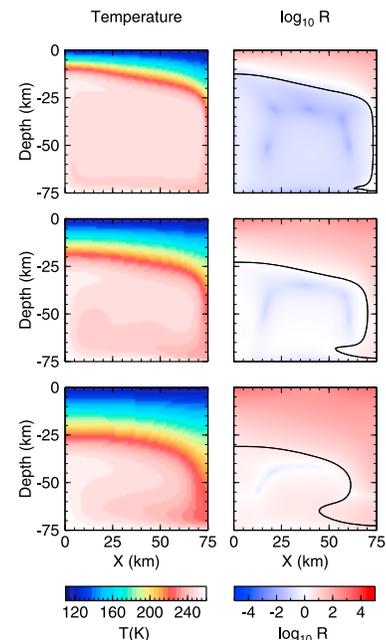
#### 4. Discussion

[14] Detection of crystal orientation fabric by radar sounding has provided valuable insights into the strain history of terrestrial ice sheets. Our simulations of solid-state convection show that COF can form and be detected in convective downwellings in an ice shell with uniform grain sizes  $d > 0.3$  mm. For larger grain sizes ( $d > 1$  mm), COF-forming regions encompass the entire ice shell. If the proposed Europa radar sounder acquires single-polarization data, only the strongest COF reflections will be detectable, but could be indistinguishable from reflections caused by changes in density, conductivity, and/or dust concentration. Multi-polarization radar could unambiguously detect COF and would require using two transmitting/receiving antennae perpendicular to each other, or flying two overlapping ground tracks with the spacecraft's radar antenna rotated  $90^\circ$  in yaw on the second pass.

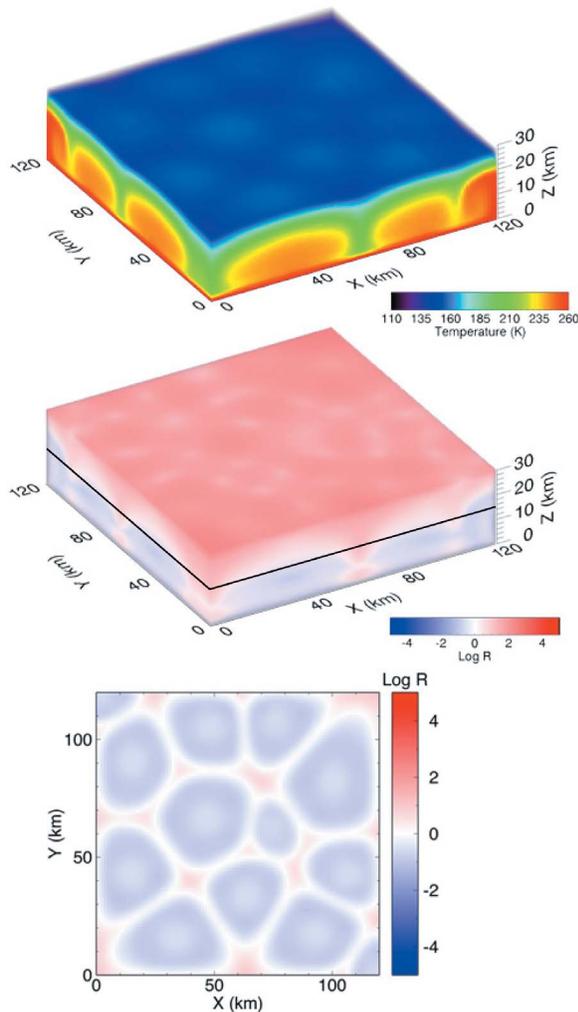
[15] A quantitative model of  $c$ -axis orientation and temperature based on the convective flow field should be meshed with a description of the locations of soluble/insoluble impurities to fully address the dielectric properties of a convecting ice shell. Such a model could be used to predict the radar returns from a realistic ice shell including polarization effects. In addition, we have used a relatively simple rheological and convective model of the satellites' ice shells. The viscosity of ice with well-developed fabric is strongly anisotropic [Cuffey and Paterson, 2010], which could lead to changes in the convection pattern and predicted strain rates in the convecting sublayer. The viscosity of ice also depends on grain size, which in turn depends on impurity content, impurity solubility, and the relative importance of dynamic recrystallization, which can occur in tandem with COF formation [e.g., Durand *et al.*, 2006; Barr and McKinnon, 2007; Cuffey and Paterson, 2010]. These effects should be included in future modeling efforts.

[16] Europa's ice shell is likely heated from within by tidal dissipation [Ojakangas and Stevenson, 1989], which can modify the convection pattern. The processes that convert tidal deformation to heat are poorly constrained [Barr and Showman, 2009], but a consensus is emerging that convective upwellings and shallow faults may be sites of intense heating [Nimmo and Gaidos, 2002; Mitri and Showman, 2008; Han and Showman, 2010]. Future work should address how tidal heating may affect the sites of COF formation within Europa's ice shell.

[17] Geological activity on Ganymede may have occurred only for a brief period [Pappalardo *et al.*, 2004], and Europa's activity may be short-lived or even cyclical



**Figure 2.** (left) Temperature and (right)  $R$  for an ice shell  $D = 75$  km thick, with a uniform grain size (top)  $d = 0.1$  mm ( $Ra_o = 2 \times 10^8$ ), (middle)  $d = 0.3$  mm ( $Ra_o = 2 \times 10^7$ ), and (bottom)  $d = 1.0$  mm ( $Ra_o = 2 \times 10^6$ ). Black line shows the location of the  $T = 240$  K isotherm. For  $T > 240$  K, COF will form if  $R > 1$  (red values). Total accumulated strain  $\epsilon \gg 1$  everywhere  $T > 240$  K.



**Figure 3.** (top) Temperature, (middle)  $R$ , and (bottom) slice of  $R$  values at the depth where the horizontally averaged  $T = 240$  K isotherm for a  $120 \times 120 \times 30$  km section of the ice I shell of a Galilean satellite with  $d = 0.3$  mm ( $Ra_o = 1.4 \times 10^6$ ). In Figures 3 (middle) and 3 (bottom), red areas within convective downwellings are locations where COF is created. Blue areas show locations where strain is accommodated by volume diffusion, which destroys COF.

[Greeley *et al.*, 2004]. Past convection could be detected by looking for a layer of relict crystal fabric within a conductive ice shell. The presence of COF beneath surface features such as european bands or Ganymede grooved terrain could test the hypotheses that these features formed from extensional strain associated with upwellings of warm ice. For example, gray bands on Europa have been suggested to form in a manner similar to terrestrial midocean ridges [Prockter *et al.*, 2002]. Modeling of their formation suggests  $\varepsilon \gg 1$ , extensional strain accommodated by dislocation creep (i.e.,  $R \gg 1$ ), and  $T \sim 260$  K only a few kilometers below the surface [Stempel *et al.*, 2005]. COF could also form beneath Ganymede's grooved terrain, where  $\varepsilon \sim 0.5$  to 2 [Pappalardo and Collins, 2005], flow in the subsurface is accommodated by dislocation creep, and  $T > 240$  K at a depth of a few kilometers [Dombard and McKinnon, 2001].

[18] **Acknowledgments.** This work is supported by the NASA OPR grant NNX09AP30G to A. C. Barr and NASA PG&G grant NNX10AJ65G to D. E. Stillman. We thank Roger Phillips and Joe MacGregor for helpful discussions.

[19] The Editor would like to thank the two anonymous reviewers for their assistance in evaluating this paper.

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